

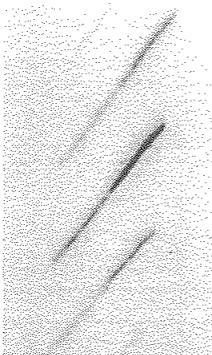


**Results** : taking advantage of the small polychromatic focus point of ID24 (80  $\mu\text{m}$  FW at 10%), the sample was easily aligned in the diffracting position, despite its very **small size** (200x100x10  $\mu\text{m}$ ). This demonstrates the feasibility of DAFS measurements on ID24 even for small crystals.

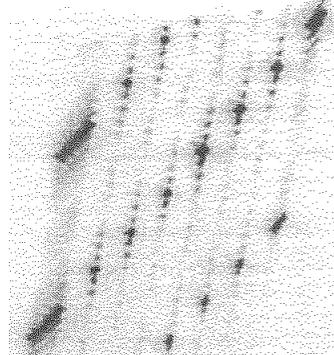
All diffraction images were collected using the rotating method, with a large 2D detector (image-plate with on-line reading). We thus collected 70 couples of images (the second one being taken with “finger slits” as energy markers, see report of exp n°MI-119 for details), both at room temperature and above the phase transition. An example of the images obtained is shown in fig. 1.

We have observed unexpected features, such as a superstructure due to stacking faults in the sample (fig. 2). This emphasizes the importance of a large **2D detector** to provide a large vision of what *actually* (expected or not) happens in the compound, which is of course very important for phase transitions.

The extraction of data from the images follows the process described for the experiment n°MI-1 19 (image corrections, integration, energy calibration, intensity normalization). But since the quantity of data (hundreds of reflections, 3 gigabytes of data) is very important, we are developing a program based on IDL language : this software will allow data reduction for all dispersive DAFS measurements in the rotating method, including indexation of reflections (main peaks or satellites ; with a refinement of experimental parameters), *automatic integration of  $I(E)$  spectra* with background subtraction, and *collection of orientation parameters* for each reflection (to allow absorption and lorentz-polarization corrections in further refinement). The development of this program is well under way and data reduction for this experiment will follow.



**Fig 1.** A row of satellite reflections. Each line correspond to a full  $I(E)$  spectra for one reflection : the energy varies linearly along the line, with longer wavelengths on the upper (right) end of the line, and shorter wavelengths on the other end. The step in intensity marks the  $W L_{III}$  edge. The energy range is  $\approx 600$  eV.



**Fig.2.** Superstructure observed in the sample, due to stacking faults (the structure of the crystal consists in slabs of 13  $WO_6$  octahedras, and some slabs have only 12). This shows a long-range order with  $\lambda \approx 600$  Å along  $c^*$ . For satellite reflections, this superstructure is replaced by diffuse scattering. (this image was taken with "slits" used as energy markers, hence the slicing of the lines)